# Prompt X-Rays from Fast-Neutron-Induced Fission of <sup>238</sup>U

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## **Fission X-Rays a Very Brief History**

- 1960-70's Reisdorf, Griffin, Wilhelmy measurements for spontaneous fission and thermal neutron induced fission only
  1965 Glendinin and Griffin calculated K x-ray yields
- Internal Conversion not fission acceleration ionization!
- Produced by levels with strong internal conversion means odd mass and odd-Z-odd-N nuclei favored
- Provide Z identification
- Complements gamma-ray studies that are most sensitive to eveneven nuclei via 2<sup>+</sup> – 0<sup>+</sup> observations
- LANSCE with CEA-Bruyeres, first measurements on <sup>238</sup>U(n,f) with fast neutrons



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# **Overview of the LANSCE/WNR Facility Showing** the Beam Structure and Neutron Flux

#### WNR Spallation Neutron Source



- Intense high-energy "white" neutron source
- $0.1 < E_n < 400 \text{ MeV}$
- Time-of-flight for efficient excitation function acquisition





1000

100

10

100

En (MeV)

**4**725 μs**→** 

**Typical WNR Proton Beam Parameters** 

8.3 ms



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1000

#### **GEANIE – Germanium Array for Neutron-Induced Excitations at LANSCE**

- Located at the WNR spallation neutron source – driven by the 800 MeV LANSCE proton linac
- Neutron energy is determined by timeof-flight on a 20 meter flight path
- Typical neutron energy range is
- 1 < En < 200 MeV
- Both 25% coaxial HPGe detectors and low-energy planar HPGe detectors are used.
- Typical gamma-ray energy range 15 keV < Eg < 4 MeV</li>
- Built on the former HERA array from Lawrence Berkeley National Laboratory







## Multiple Solar Cell Fission Fragment Detectors Were Used with <sup>238</sup>U Deposited by Mass Separator (CEA)

- 11 low energy photon spectrometers for x-ray detection
- 15 coaxial Ge detectors for γ ray detection
- 8 <sup>238</sup>U deposits on thin solar cells in the WNR neutron beam as an active target
- Fission-photon coincidences required to eliminate high backgrounds at Eγ<50 keV</li>



Si solar cells (< 100  $\mu$ m)

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U deposits (~1 mg/cm<sup>2</sup>)
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Neutron Beam

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#### **X-Ray Spectra for Lighter-Mass Fission Fragments**





#### **X-Ray Spectra for Higher-Mass Fission Fragments**



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## **Fission Fragment K X-Ray Decay Properties**

- For our data, measured on 100 ns time scale, <u>mainly E2</u>, <u>M1 transitions</u> contribute significantly to the observed Kshell x-ray yields
- Internal conversion rate is small for E1 transitions
- Decay time is much longer for higher multipoles and conversion is much less for higher gamma energies
- Odd-even staggering is observed, especially for heavier fragments





## <sup>238</sup>U(n,f) Mass Yield Distributions vs E<sub>n</sub> (for lodine) Provide a View into Different Masses with changing E<sub>n</sub>





# Wahl Fission Fragment Systematics Compared with GEANIE Gamma-Ray Yield Data for <sup>238</sup>U(n,f) vs E<sub>n</sub>



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https://www-nds.iaea.org/publications/tecdocs/sti-pub-1286/



#### Measured K X-Ray Yields vs Atomic Number for Five Incident Neutron Energy Bins from 3 to 180 MeV



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## Calculations of X-Ray Yields from Energy Levels, Fission Yield Systematics, and other data

- Calculated K Yield = Sum over all known (NUDAT) IC levels (α/(1+α) weighted by systematic (Wahl) mass yields \* branching \* lifetime factor \* fluorescent yield
- Consider the case of a single low-lying state with large internal conversion coefficient
  - Typically have large feeding from higher levels
  - For ease of calculations assume 100% population
  - But, may be less due to isomers, feeding patterns
  - Multiple x-ray emission is possible
- Estimated uncertainties in calculations include only IC coefficient and mass yield uncertainties



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#### **E**<sub>n</sub>=3 MeV Measured and Calculated K X-Ray Yields

- Data greater than calculation missing IC levels
- Data less than calculation 100% feeding not true





#### **E**<sub>n</sub>=8 MeV Measured and Calculated K X-Ray Yields









#### **E**<sub>n</sub>=14 MeV Measured and Calculated K X-Ray Yields





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#### **E**<sub>n</sub>=32 MeV Measured and Calculated K X-Ray Yields





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#### **E**<sub>n</sub>=180 MeV Measured and Calculated K X-Ray Yields





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## **Reisdorf <sup>252</sup>Cf Spontaneous Fission K Yields**







## K X-ray yields for thermal fission of <sup>233,235</sup>U, <sup>239</sup>Pu, & spontaneous fission of <sup>252</sup>Cf



Fig. Top: Mean K X-ray emission probabilities per fragment of charge Z (K(Z)) obtained by Reisdorf *et al.* [4] for <sup>252</sup>Cf spontaneous fission (circles), <sup>235</sup>U(n,f) (squares), <sup>233</sup>U(n,f) (triangles) and <sup>239</sup>Pu(n,f) (diamonds). Bottom: Ratio of K(Z) in <sup>252</sup>Cf spontaneous fission (circles), <sup>233</sup>U(n,f) (triangles) and <sup>239</sup>Pu(n,f) (diamonds) to K(Z) for <sup>235</sup>U (data from ref. [4]).







## **Charge yields inferred from K X-ray yields**





## Inferred charge yields for 180 MeV



Fig. Charge distributions determined from the X-ray yield measurements (symbols). Top: threshold–6 MeV,  $\langle E_n \rangle \simeq$  3 MeV. Bottom: 6–11 MeV,  $\langle E_n \rangle \simeq$  8 MeV. Solid error bars correspond to propagated fit errors, dotted error bars correspond to the sum in quadrature of the fit error and the 50% relative uncertainty associated with K(Z). Dashed curve is the result of the GEF code by Schmidt-Jurado [43] (see text). Solid curve within hatched area corresponds to Wahl systematics and associated uncertainty obtained for the corresponding energy ranges.

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## Future

- High-energy neutron-induced fission of actinides provides a window for spectroscopy of a range of neutron-rich nuclei
- X-ray gamma-ray coincidences can provide new information on energy levels of specific neutron rich nuclei
- Fission X-ray-gamma-ray coincidence experiments have been proposed using increased actinide mass in a fission counter to enable x-ray-gamma-ray coincidence studies with sufficient statistics



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